

1743 - Constraining the Atmosphere of the Terrestrial Exoplanet Gl486b

Cycle: 1, Proposal Category: GO

INVESTIGATORS

Name	Institution	E-Mail
Megan Mansfield (PI)	University of Arizona	meganmansfield@uchicago.edu
Prof. Jacob L. Bean (CoI)	University of Chicago	jbean@astro.uchicago.edu
Prof. Eliza MR. Kempton (CoI)	University of Maryland	ekempton@astro.umd.edu
Prof. Edwin S Kite (CoI)	University of Chicago	kite@uchicago.edu
Daniel Koll (CoI)	Massachusetts Institute of Technology	dkoll@mit.edu
Matej Malik (CoI)	University of Maryland	malik@umd.edu

OBSERVATIONS

Folder	Observation	Label	Observing Template	Science Target
Gl 486t)			
	1	Gl486_visit1	MIRI Low Resolution Spectroscopy	(1) GL486
	2	Gl486_visit2	MIRI Low Resolution Spectroscopy	(1) GL486

ABSTRACT

JWST will provide the first opportunity to search for signs of habitability on terrestrial planets orbiting M dwarf stars. However, it is currently unknown whether planets around M dwarfs can retain atmospheres given the relatively violent history of extreme UV irradiation from their host stars. The first step to understanding the habitability of rocky planets is therefore to establish whether M dwarf terrestrial planets can retain significant atmospheres. Furthermore, terrestrial exoplanets well inside the inner edge of the habitable zone offer a chance to test theories of atmospheric evolution in a previously unexplored regime.

We propose to use MIRI/LRS secondary eclipse observations to search for the presence of an atmosphere on the terrestrial M dwarf planet Gl 486b (TOI 1827.01) and to constrain its composition. Gl 486b is the highest signal-to-noise terrestrial planet for these observations other than LHS 3844b,

JWST Proposal 1743 (Created: Thursday, July 22, 2021 at 12:00:48 PM Eastern Standard Time) - Overview

which has been previously observed with Spitzer and found to not posess an atmosphere. Additionally, Gl 486b has less than one quarter of the integrated XUV flux of LHS 3844b, which means it is much more likely to have retained an atmosphere. We find that with two secondary eclipse observations of Gl 486b we can determine whether or not it has an atmosphere with a surface pressure of P1 bar at the 5-sigma significance level. The observed spectrum will also constrain the composition of its atmosphere (if one is observed) or surface (if no atmosphere is detected) at the 3-sigma significance level. These observations will provide valuable data for constraining the impact of M dwarf irradiation on terrestrial planet atmospheres.

OBSERVING DESCRIPTION

We will perform time-series observations to measure the secondary eclipse of the exoplanet Gl 486b. The fundamental parameter that determines the length of our time-series observations is the eclipse duration of Gl 486b, which is 1.08 hours. In addition to this length of time, we need to observe out-of-eclipse baseline in order to identify and effectively model any instrument systematics. Such systematics typically introduce red noise, which can occur on varying timescales and requires longer observational baselines to identify, characterize, and remove. We include 3 hours of baseline, as a 3/1 ratio of baseline to in-eclipse observing should allow us to successfully identify and remove systematics. We also include 30 minutes to allow the telescope to settle. Including such a settling time has become standard in space-based exoplanet observations because of the increased influence of instrument systematics immediately after repointing. For example, Spitzer Space Telescope observations often included 30 minutes of settling time and Hubble Space Telescope observations frequently included an orbit (about 45 minutes) of settling time. Including overheads, the total time we request for these observations is therefore 6.3 hours per eclipse, or a total of 12.6 hours for the two eclipse observations.

MIRI/LRS time-series observations use the SLITLESSPRISM mode without dithers. We used PandExo to simulate our proposed observations and determine the optimal observing pattern. We will use the FAST readout pattern with 5 groups/integration and 20735 integrations/exposure, for a total of 4.58 hours of observing time per visit. This observing strategy will ensure we remain below 76% of the saturation limit and has an overall efficiency of 80%. The visit must cover the secondary eclipse of Gl 486b, so the phase constraints we have placed ensure that there will be at least 1 hour of out-of-eclipse baseline on either side of the eclipse.

We will use the F1500W filter for target acquisition. Our single target acquisition integration will consist of 4 groups, which will result in a SNR of 420 for target acquisition while staying at 17% of the saturation limit.

Proposal 1743 - Targets - Constraining the Atmosphere of the Terrestrial Exoplanet Gl486b

	#	Name	Target Coordinates	Targ. Coord. Corrections	Miscellaneous
its	(1)	GL486	RA: 12 47 55.5675 (191.9815313d)	Proper Motion RA: -0.06822546051816196 sec of	
rget			Dec: +09 44 57.91 (9.74942d)	time/yr	
Ta			Equinox: J2000	Proper Motion Dec: -0.45979999993051024 arcsec/yr	
				Epoch of Position: 2015.5	
×			argetselector and retrieved from the SIMBAD database.		
	Category=Sta Description=				
	Extended=NO				

Pro	posal 1743 - Observation	n 1 - Constraining	the Atmosphe	re of the Terre	estrial Exoplanet	GI486b					
Observation											
Diagnostics	(Gl486_visit1 (Obs 1)) Warning (Form) (Visit 1:1) Warning (Form): Overheads	-		onds. Above this limit	it is possible that a High Ga	ain Antenna move may o	ccur during the exposu	re.			
	# Name	Target Coordinates		U	oord. Corrections		scellaneous				
Targets	(1) GL486	RA: 12 47 55.5675 (· /	Proper M time/yr	Motion RA: -0.0682254605	1816196 sec of					
arg		Dec: +09 44 57.91 (9	9.74942d)	•	Motion Dec: -0.4597999999	03051024 arcsec/yr					
ΗË		Equinox: J2000		-	f Position: 2015.5	555162 (alesses y)					
Fixed	Comments: This object was generated b Category=Star Description=[Exoplanets] Extended=NO	ny the targetselector and retr	ieved from the SIMBAD a	-							
u	# Target	Filter	Readout Pattern	Groups/Int	Integrations/Exp	Total Integrations	Total Exposure Ti	ne ETC Wkbk.Calc ID			
Acquisition	1 SAME	F1500W	FAST	4	1	1	0.636	59209.2			
ite	Subarray			Obtai	n Verification Image?						
Template	SLITLESSPRISM			true							
rs	# Dit	her Type	No. Spectral Steps	Spect	ral Step Offset	No. Spatial Steps	Spatial	Step Offset			
Dithers	1 NC	DNE									
u	# Readou	t Pattern Grou	ıps/Int	Integrations/Exp	Total Integration	ns Total Expo	osure Time El	C Wkbk.Calc ID			
Pointing Verification	1 FAST	4		1	1	0.636		209.2			

#	Readout Pattern	Groups/Int	Integrations/Exp	Total Integrations	Exposures/Dith	Total Dithers	Total Exposure Time	ETC Wkbk.Calc ID
Spectral Elements	FASTR1	6	14810	14810	1	1	16487.518	82457.1
	0.456 with period 1.46713 Da servation	ys and zero-phase 24:	58931.1593 HJD					

Pro	<u>posal 1743 - Obser</u>	vation 2 - Constra	aining the Atmosphe	re of the Terr	estrial Exoplanet	GI486b					
Observation											
Diagnostics		-	n exceeds the limit of 10000.0 sec l the Visit Planner has been run.	onds. Above this limit	: it is possible that a High Ga	ain Antenna move may o	occur during the exposure	<u>.</u>			
	# Name	Target Coo		0	Coord. Corrections		iscellaneous				
Targets	(1) GL486		55.5675 (191.9815313d)	Proper time/yr	Motion RA: -0.0682254605	1816196 sec of					
arg			4 57.91 (9.74942d)	•	Motion Dec: -0.4597999999	03051024 arcsec/vr					
Ľ		Equinox: J2	2000	-	of Position: 2015.5	20021021 alesses, yr					
Fixed	Comments: This object was gen Category=Star Description=[Exoplanets] Extended=NO	erated by the targetselector	and retrieved from the SIMBAD o	latabase.							
u u	# Target	Filter	Readout Pattern	Groups/Int	Integrations/Exp	Total Integrations	Total Exposure Tim	e ETC Wkbk.Calc ID			
Acquisition	1 SAME	F1500W	FAST	4	1	1	0.636	59209.2			
te	Subarray	Subarray Obtain Verification Image?									
Template	SLITLESSPRISM			true							
rs	#	Dither Type	No. Spectral Steps	Spec	tral Step Offset	No. Spatial Steps	Spatial S	tep Offset			
Dithers	1	NONE									
۲ 2	#	Readout Pattern	Groups/Int	Integrations/Exp	Total Integration	ns Total Expo	osure Time ET	C Wkbk.Calc ID			
Pointing Verification	1	FAST	4	1	1	0.636	592	09.2			

ţs	#	Readout Pattern	Groups/Int	Integrations/Exp	Total Integrations	Exposures/Dith	Total Dithers	Total Exposure Time	ETC Wkbk.Calc ID
Spectral Elements	1	FASTR1	6	14810	14810	1	1	16487.518	82457.1
Special Requirements	Phase 0.414 to 0.456 w Time Series Observatio No Parallel	rith period 1.46713 Day	s and zero-phase 24589	31.1593 HJD					